

## THE ACCURATE FOURTH-ORDER DOPPLER PARAMETER CALCULATION AND ANALYSIS FOR GEOSYNCHRONOUS SAR

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**Abstract**—This paper presents a new approach to calculate the 1–4th order Doppler parameters for Geosynchronous Synthetic Aperture Radar (Geo-SAR). To get accurate calculation results, the Earth is modeled as an ellipsoid and the relative motion between the sensor in an elliptical orbit and the rotating Earth is analyzed. The  $J_2$ ,  $J_3$  and  $J_4$  orbital perturbation items and attitude steering are analyzed. Ignoring the perturbation force would produce errors of the Doppler parameters for spaceborne SAR because it can influence the six orbital elements. Since the Doppler parameters are related to the antenna beam pointing directions and influenced by attitude of SAR platform, the calculation results before and after attitude steering are shown. Furthermore, the Doppler parameter properties during the whole orbital periods of Geo-SAR are compared with those of Low-Earth-Orbital SAR (Leo-SAR). Finally, the effects on Doppler parameters stemmed from the radar beam pointing accuracy are analyzed.

### 1. INTRODUCTION

Currently, most of the spaceborne SAR systems are considered as Leo-SAR since they fly at altitudes of 560–870 km. Observation of Earth surface needs large swath widths and short revisit cycle, however, Leo-SAR systems are limited by the low flying altitudes. The Geo-SAR, proposed by Tomiyasu et al., exhibits distinctive advantages which

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provide an enormous instantaneous field of view with orbital period equal to one Earth day [1–9].

A main problem of it is the super-long integration time to traverse the long non-linear synthetic aperture distance. For example, it needs up to 23 min in some conditions. This is quite different from Leo-SAR. Since a more accurate range model and a different azimuth compression reference function are required for generating high quality images, the 1–4th order Doppler parameters are necessary. Various Doppler parameter calculation algorithms have been reported for Leo-SAR, but show large errors when applied to Geo-SAR due to geometrical approximation [10–13]. Actually, in the satellite case, the geometry is more complicated, as the orbit is elliptical and the Earth is modeled as an ellipsoid rotating independently of the sensor trajectory. An accurate approach of calculating the 1–4th order Doppler parameters in Geo-SAR, which utilizes the data such as the spacecraft ephemeris and attitude, is proposed in this paper.

Furthermore, we analyze the error of Doppler parameter calculation results caused by ignoring orbital perturbation. Due to a variety of perturbing effects, the so-called *six orbital elements* describing the orbit increase or decrease with time, and the satellite orbit could not be considered as an ellipse strictly in an invariant plane. As a result, the state vectors of sensor and target will be different from those in a strict elliptical orbit [15]. As a result, the Doppler parameters, without considering the perturbation force, will not be accurate enough for precision focusing for the spaceborne SAR data. In Leo-SAR condition, the influence from the Earth shape perturbation is the main part. As a result, we analyze the varieties of the orbital elements under orbital perturbative force when the terms  $J_2$ ,  $J_3$  and  $J_4$  are all considered, and the values of Doppler parameters are analyzed.

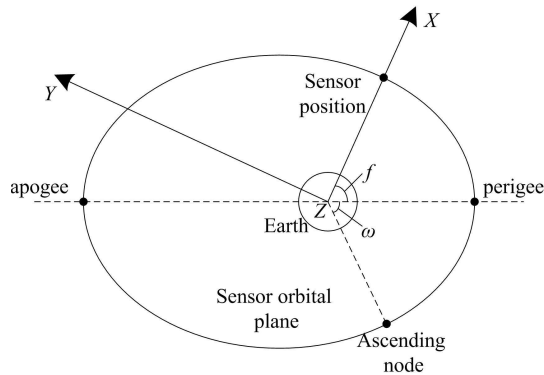
It should be emphasized that the attitude steering is necessary for Geo-SAR. Since the Doppler parameter calculation results depend on the beam pointing considerably, the effects of antenna yaw and pitch on the Doppler parameters are simulated and analyzed. Finally, the beam pointing errors of  $0.05^\circ$  for yaw, pitch and roll angles are assumed, and the effects of pointing errors on the 1–4th order Doppler parameters are illustrated.

This paper is structured as follows. In Section 2, a 3D geometry model is illustrated, and the state vectors of sensor and target are derived based on the model. The 1–4th order Doppler parameter calculation expressions are then advanced. The orbital perturbation is analyzed in Section 3. In Section 4, we discuss the attitude steering. The simulation results are shown in Section 5. Finally, the conclusion is given in Section 6.

## 2. FOURTH-ORDER DOPPLER PARAMETERS CALCULATION

### 2.1. Accurate Geometry Model

A satellite's orbit can be described by six orbital elements, which are constants of the ellipse and of its orientation relative to the Earth Centered Inertial coordinate system (ECI). They are:  $a$ , the semi-major axis of the orbit;  $e$ , the eccentricity of the orbit;  $\theta_i$ , the inclination of the orbit;  $\Omega$ , the longitude of the ascending node;  $\omega$ , the argument of perigee;  $M$ , the mean anomaly. As depicted in Fig. 1, consider a satellite local coordinate (SLC) system with origin at the center of the Earth. The  $X$  axis of SLC always points to the sensor, the  $Z$  axis is perpendicular to the orbital plane, and the  $Y$  axis obeys the Cartesian right hand rule. The angle  $f$  is true anomaly.



**Figure 1.** Geo-SAR geometry model in SLC.

To get the 1–4th order Doppler parameters, the 1–5th order state vectors of sensor and target should be presented. The symbols  $\mathbf{r}$ ,  $\mathbf{v}$ ,  $\mathbf{A}$ ,  $\mathbf{A}'$  and  $\mathbf{A}''$  are respectively the state vectors of position, velocity, acceleration, rate of acceleration, and the second derivative of acceleration. The five state vectors of Geo-SAR can be separately expressed as follows, where the subscript  $s$  means satellite:

$$\mathbf{r}_s = (R_s, 0, 0)^T \tag{1}$$

$$\mathbf{v}_s = \sqrt{\mu/[a(1-e^2)]} (e \sin f, 1 + e \cos f, 0)^T \tag{2}$$

$$\mathbf{A}_s = (-\mu/R_s^2, 0, 0)^T \tag{3}$$

$$\mathbf{A}'_s = (2\mu R'_s/R_s^3, -\mu\omega_s/R_s^2, 0)^T \tag{4}$$

$$\mathbf{A}_s'' = \left( \frac{2\mu R_s'' R_s - 6\mu R_s'^2 + \mu \omega_s^2 R_s^2}{R_s^4}, \frac{4\mu R_s' \omega_s - \mu \omega_s' R_s}{R_s^3}, 0 \right)^T \quad (5)$$

where  $\mu$  is the gravitational constant,  $R_s$  instantaneous distances between sensor and the Earth center,  $R_s'$  the first-order temporal derivative of  $R_s$ ,  $R_s''$  the second-order temporal derivative of  $R_s$ ,  $\omega_s$  instantaneous angular velocity of sensor,  $\omega_s'$  the first-order temporal derivative of  $\omega_s$ . According to Kepler's second law,  $R_s$ ,  $R_s'$ ,  $R_s''$ ,  $\omega_s$ , and  $\omega_s'$  all change with the true anomaly during the whole orbital periods [2], and they can be respectively presented as:

$$R_s = [a(1 - e^2)] / (1 + e \cos f) \quad (6)$$

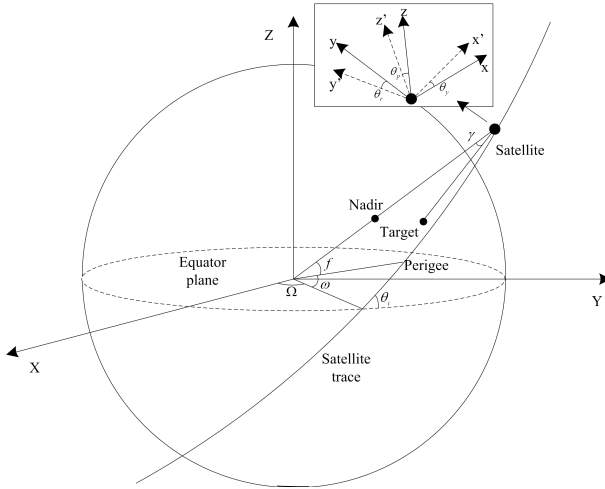
$$R_s' = e \sin f \sqrt{\mu / [a(1 - e^2)]} \quad (7)$$

$$R_s'' = \mu e \cos f (1 + e \cos f)^2 / [a^2 (1 - e^2)^2] \quad (8)$$

$$\omega_s = \sqrt{\mu / [a(1 - e^2)]^3} (1 + e \cos f)^2 \quad (9)$$

$$\omega_s' = -2\mu e \sin f (1 + e \cos f)^3 / [a^3 (1 - e^2)^3] \quad (10)$$

When the satellite moves in the orbit, the radar beam center line intersects with the Earth surface to form a group of points, known as targets. As depicted in Fig. 2, the five state vectors of the targets are presented as follows, where the subscript  $t$  means targets.



**Figure 2.** Geo-SAR geometry in 3D space.

$$\mathbf{r}_t = A_{re}A_{ea}(-r, 0, 0)^T + \mathbf{r}_s \quad (11)$$

$$\mathbf{v}_t = A_{rv}A_{vo} \begin{pmatrix} -\omega_e R_a \cos \theta_{lat} \sin \theta_{long} \\ \omega_e R_a \cos \theta_{lat} \cos \theta_{long} \\ 0 \end{pmatrix} \quad (12)$$

$$\mathbf{A}_t = A_{rv}A_{vo} \begin{pmatrix} -\omega_e^2 R_a \cos \theta_{lat} \cos \theta_{long} \\ -\omega_e^2 R_a \cos \theta_{lat} \sin \theta_{long} \\ 0 \end{pmatrix} \quad (13)$$

$$\mathbf{A}'_t = A_{rv}A_{vo} \begin{pmatrix} \omega_e^3 R_a \cos \theta_{lat} \sin \theta_{long} \\ -\omega_e^3 R_a \cos \theta_{lat} \cos \theta_{long} \\ 0 \end{pmatrix} \quad (14)$$

$$\mathbf{A}''_t = A_{rv}A_{vo} \begin{pmatrix} \omega_e^4 R_a \cos \theta_{lat} \cos \theta_{long} \\ \omega_e^4 R_a \cos \theta_{lat} \sin \theta_{long} \\ 0 \end{pmatrix} \quad (15)$$

where  $A_{ea}$ ,  $A_{re}$ ,  $A_{rv}$  and  $A_{vo}$  are the Euler transformation matrixes:

$$A_{ea} = \begin{pmatrix} \cos \gamma & 0 & -k \sin \gamma \\ 0 & 1 & 0 \\ k \sin \gamma & 0 & \cos \gamma \end{pmatrix} \quad (16)$$

$$A_{re} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_y & \sin \theta_y \\ 0 & -\sin \theta_y & \cos \theta_y \end{pmatrix} \begin{pmatrix} \cos \theta_p & \sin \theta_p & 0 \\ -\sin \theta_p & \cos \theta_p & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos \theta_r & 0 & -\sin \theta_r \\ 0 & 1 & 0 \\ \sin \theta_r & 0 & \cos \theta_r \end{pmatrix} \quad (17)$$

$$A_{rv} = \begin{pmatrix} \cos f & \sin f & 0 \\ -\sin f & \cos f & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad (18)$$

$$A_{vo} = \begin{pmatrix} \cos \omega & \sin \omega & 0 \\ -\sin \omega & \cos \omega & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_i & \sin \theta_i \\ 0 & -\sin \theta_i & \cos \theta_i \end{pmatrix} \begin{pmatrix} \cos \Omega & \sin \Omega & 0 \\ -\sin \Omega & \cos \Omega & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad (19)$$

where  $\omega_e$  is rotational angular velocity of the Earth,  $\gamma$  the elevation angle,  $\theta_y$  the yaw angle,  $\theta_p$  the pitch angle, and  $\theta_r$  the roll angle, and  $k$  the beam pointing direction ( $k = 1$  for right look, and  $k = -1$  for left look). The value of  $r$  is the instantaneous range from sensor to targets,  $R_a$  the local Earth radii of the targets, and  $\theta_{lat}$  and  $\theta_{long}$  are respectively local latitudes and longitudes of the targets. The expressions of  $r$ ,  $R_a$ ,  $\theta_{lat}$  and  $\theta_{long}$  can be derived in the Earth centered inertial coordinates (ECI) directly. According to the discussions above, the sensor's coordinates in ECI is:

$$\begin{pmatrix} x_s \\ y_s \\ z_s \end{pmatrix} = \begin{pmatrix} R_s \cos(f + \omega) \\ R_s \sin(f + \omega) \cos \theta_i \\ R_s \sin(f + \omega) \sin \theta_i \end{pmatrix} \quad (20)$$

After a series of derivations, the coordinates of the targets in ECI can be expressed as:

$$\begin{pmatrix} X_t \\ Y_t \\ Z_t \end{pmatrix} = \begin{pmatrix} Ar + x_s \\ Br + y_s \\ Cr + z_s \end{pmatrix} \quad (21)$$

where  $A$ ,  $B$  and  $C$  can be respectively given as:

$$\begin{aligned} A = & -\cos(f + \omega) \cos \theta_p \cos(\gamma + \theta_r) \\ & + \sin(f + \omega) \cos \theta_y \sin \theta_p \cos(\gamma + \theta_r) \\ & + k \sin(f + \omega) \sin \theta_y \sin(\gamma + \theta_r) \end{aligned} \quad (22)$$

$$\begin{aligned} B = & -\cos \theta_i \sin(f + \omega) \cos \theta_p \cos(\gamma + \theta_r) \\ & - \cos \theta_i \cos(f + \omega) \cos \theta_y \sin \theta_p \cos(\gamma + \theta_r) \\ & - k \cos \theta_i \cos(f + \omega) \sin \theta_y \sin(\gamma + \theta_r) \\ & - \sin \theta_i \sin \theta_y \sin \theta_p \cos(\gamma + \theta_r) + k \sin \theta_i \cos \theta_y \sin(\gamma + \theta_r) \end{aligned} \quad (23)$$

$$\begin{aligned} C = & -\sin \theta_i \sin(f + \omega) \cos \theta_p \cos(\gamma + \theta_r) \\ & - \sin \theta_i \cos(f + \omega) \cos \theta_y \sin \theta_p \cos(\gamma + \theta_r) \\ & - k \sin \theta_i \cos(f + \omega) \sin \theta_y \sin(\gamma + \theta_r) \\ & + \cos \theta_i \sin \theta_y \sin \theta_p \cos(\gamma + \theta_r) - k \cos \theta_i \cos \theta_y \sin(\gamma + \theta_r) \end{aligned} \quad (24)$$

As the Earth surface is modeled as an ellipsoid:

$$\frac{X_t^2}{R_e^2} + \frac{Y_t^2}{R_e^2} + \frac{Z_t^2}{R_p^2} = 1 \quad (25)$$

where  $R_e = 6378.137$  km and  $R_p = 6356.752$  km are the equatorial and polar radii of the ellipsoidal Earth [1]. Substituting (21)–(24) into (25) and simplifying,  $r$  can be calculated as:

$$r = \frac{-K_2 - \sqrt{K_2^2 - 4K_1K_3}}{2K_1} \quad (26)$$

where  $K_1$ ,  $K_2$  and  $K_3$  can be given as:

$$\begin{cases} K_1 = R_p^2 A^2 + R_p^2 B^2 + R_e^2 C^2 \\ K_2 = 2(Ax_s R_p^2 + By_s R_p^2 + Cz_s R_e^2) \\ K_3 = R_p^2 x_s^2 + R_p^2 y_s^2 + R_e^2 z_s^2 - R_e^2 R_p^2 \end{cases}$$

Finally, the local latitudes, longitudes and Earth radii of the targets can be separately calculated as:

$$\theta_{lat} = \arctan \left[ \frac{Z_t}{\sqrt{X_t^2 + Y_t^2}} \right] \quad (27)$$

$$\theta_{long} = \Omega + \begin{cases} \arctan[Y_t/X_t], & (X_t > 0, Y_t > 0) \\ \arctan[Y_t/X_t] + 2\pi, & (X_t > 0, Y_t < 0) \\ \arctan[Y_t/X_t] + \pi, & (X_t < 0) \end{cases} \quad (28)$$

$$R_a = \sqrt{\frac{R_e^2 R_p^2}{[R_p \cos \theta_{lat}]^2 + [R_e \sin \theta_{lat}]^2}} \quad (29)$$

Substituting (26)–(29) into (11)–(15), the five state vectors of the targets can be obtained.

### 2.2. Doppler Parameter Calculation

This section gives a rigorous derivation of Doppler parameter expressions based on the geometry model above. The range vector from the target to the sensor in SLC can be given as:

$$\mathbf{r} = \mathbf{r}_s - \mathbf{r}_t \quad (30)$$

The expression (30) may be rewritten as:

$$r^2 = (\mathbf{r}_s - \mathbf{r}_t) (\mathbf{r}_s - \mathbf{r}_t) \quad (31)$$

Differentiating both sides of (31) with respect to time from the first-order to the fourth-order, we can get  $r'$ ,  $r''$ ,  $r^{(3)}$  and  $r^{(4)}$ . The 1–4th order Doppler parameters can then be presented as:

$$f_{dc} = -\frac{2r'}{\lambda} = -\frac{2}{\lambda} \frac{(\mathbf{r}_s - \mathbf{r}_t) (\mathbf{v}_s - \mathbf{v}_t)}{r} \quad (32)$$

$$f_{1r} = -\frac{2r''}{\lambda} = -\frac{2}{\lambda} \left[ \frac{(\mathbf{v}_s - \mathbf{v}_t)^2}{r} + \frac{(\mathbf{A}_s - \mathbf{A}_t) (\mathbf{r}_s - \mathbf{r}_t)}{r} - \frac{\lambda^2 f_{dc}^2}{4r} \right] \quad (33)$$

$$f_{2r} = -\frac{2r^{(3)}}{\lambda} = -\frac{2}{\lambda} \left[ \frac{3(\mathbf{v}_s - \mathbf{v}_t) (\mathbf{A}_s - \mathbf{A}_t)}{r} + \frac{(\mathbf{r}_s - \mathbf{r}_t) (\mathbf{A}'_s - \mathbf{A}'_t)}{r} - \frac{3\lambda^2 f_{dc} f_{1r}}{4r} \right] \quad (34)$$

$$f_{3r} = -\frac{2r^{(4)}}{\lambda} = -\frac{2}{\lambda} \left[ \frac{3(\mathbf{A}_s - \mathbf{A}_t)^2}{r} + \frac{4(\mathbf{v}_s - \mathbf{v}_t) (\mathbf{A}'_s - \mathbf{A}'_t)}{r} + \frac{(\mathbf{r}_s - \mathbf{r}_t) (\mathbf{A}''_s - \mathbf{A}''_t)}{r} - \frac{3\lambda^2 f_{1r}^2}{4r} - \frac{\lambda^2 f_{dc} f_{2r}}{r} \right] \quad (35)$$

where  $\lambda$  is the radar wavelength. Substituting (1)–(5) and (11)–(15) into (32)–(35), the 1–4th order Doppler parameters can be calculated.

### 3. PERTURBATION OF ORBIT

Since the Earth is an oblate ellipsoid, bulging at the equator, with even higher order non-sphericity, the gravitation field on the sensor

can never be considered as a central one. As a result, the six orbital elements are affected by perturbation force, and the satellite trajectory is not a simple ellipse in an invariant plane. The  $J_2$ ,  $J_3$  and  $J_4$  perturbation terms and their effect on the Doppler parameters are analyzed in this paper. The gravitational potential function of the Earth is written as [15]:

$$U = \frac{\mu}{r} \left[ 1 - \frac{J_2 R_e^2}{2R_s^2} (3 \sin^2 \varphi - 1) - \frac{J_3 R_e^3}{2R_s^3} (5 \sin^3 \varphi - 3 \sin \varphi) - \frac{J_4 R_e^4}{8R_s^4} (35 \sin^4 \varphi - 30 \sin^2 \varphi + 3) \right] \quad (36)$$

where  $\varphi$  is the local latitude of nadir, and it is convenient to be expressed as follow:

$$\varphi = \arcsin [\sin \theta_i \sin(f + \omega)] \quad (37)$$

According to (36), the perturbation potential function due to the Earth's oblateness is given by:

$$U_P = U - \frac{\mu}{r} \quad (38)$$

The perturbation force of satellite at the corresponding orbital position can then be represented as:

$$\mathbf{F} = \text{grad}(U_P) \quad (39)$$

Its influence upon the satellite orbit is illuminated by Kozai, and the periodic perturbations of the first order and secular perturbations up to the second order are derived. The symbol  $\chi$  represents the six orbital elements ( $a$ ,  $e$ ,  $\theta_i$ ,  $\Omega$ ,  $\omega$  and  $M$ ). Based on the conclusion above, the expressions of the six orbital elements under perturbation force have the same form as Equation (40).

$$\chi_P(t) = \overline{\chi_0} + \chi_1(t - t_0) + \chi_2(t - t_0) + \chi_s^{(1)}(t) + \chi_L^{(1)}(t) \quad (40)$$

where  $\chi_P(t)$  is the orbital element under perturbation force,  $\overline{\chi_0}$  the mean orbital element value,  $\chi_1(t - t_0)$  the first-order secular term,  $\chi_2(t - t_0)$  the second-order secular term,  $\chi_s^{(1)}(t)$  the first-order short periodic term,  $\chi_L^{(1)}(t)$  the first-order long periodic term,  $t$  the orbital time, and  $t_0$  the start time. As a result, the actual six orbital elements, which are under perturbation force, are denoted by  $a_P(t)$ ,  $e_P(t)$ ,  $\theta_{iP}(t)$ ,  $\Omega_P(t)$ ,  $\omega_P(t)$  and  $M_P(t)$  respectively.

According to the theory in Section 2, the state vectors of sensor and target are decided by the six orbit elements and radar's beam pointing direction. During one orbital period of the sensor, the accumulated changes of the state vectors due to the varieties of six



orbital elements can never be neglected. Assuming that the beam pointing direction is invariant, the state vectors depend on the six orbital elements only. The expression of the 1–4th order Doppler parameter of spaceborne SAR is given by:

$$f_{nrP} = \Psi_n [a_P(t), e_P(t), \theta_{iP}(t), \omega_P(t), \Omega_P(t), M_P(t)], \quad (n = 0, 1, 2, 3) \quad (41)$$

where  $f_{nrP}$  is the  $n$ th-order Doppler parameter ( $f_{0rP}$  means Doppler centroid,  $f_{dc}$ ), the subscript  $P$  means “under perturbation force”, and the symbol  $\Psi_n$  represents “function of some parameters”. The error in percentage of  $f_{nr}$  can be calculated by:

$$E_{f_{nr}} = \frac{(f_{nrP} - f_{nr})}{f_{nrP}} \times 100, \quad (n = 0, 1, 2, 3) \quad (42)$$

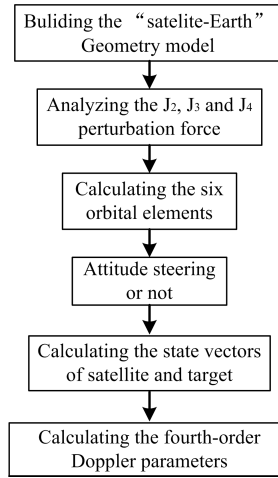
The results will be simulated and analyzed in Section 5.

#### 4. ATTITUDE STEERING

In addition to the orbital perturbation, the beam pointing also has an obvious effect on the Doppler parameters. Due to the Earth rotation and orbital eccentricity, the periodic yaw and pitch angle are generated. The attitude steering should be performed to get the zero Doppler centroid over the entire swath during the whole orbital periods. The approach proposed in [16] has a good effect on Geo-SAR. After attitude steering has been applied, the beam center line direction will be quite different, and the Doppler parameters change significantly. A block diagram of steps to calculate the spaceborne SAR’s Doppler parameters is shown in Fig. 3.

#### 5. SIMULATION RESULTS

The Doppler parameter calculation error caused by ignoring orbital perturbation is illustrated in Section 5.1. The simulation and comparison of 1–4th order Doppler parameters before and after attitude steering are described in Section 5.2. The effects on Doppler parameters caused by attitude control error are presented in Section 5.3. Since the Doppler parameter properties of Geo-SAR are quite different from those of Leo-SAR, the comparison between them is shown in this section. Without loss of generality, TerraSAR-X is chosen as an example of Leo-SAR [14]. The simulation parameters are listed in Table 1.



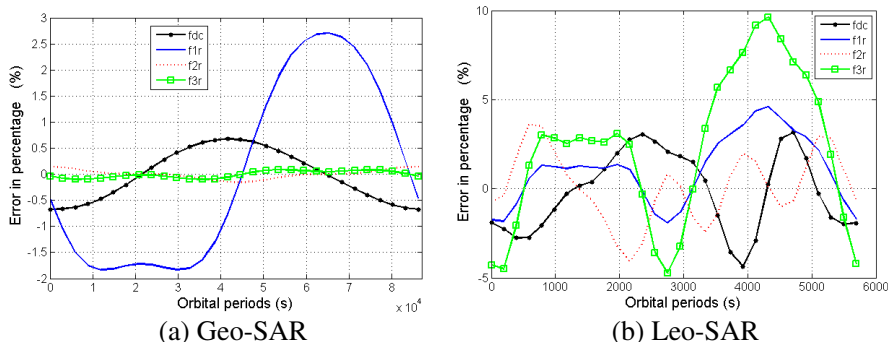
**Figure 3.** Doppler parameter calculation block diagram of spaceborne SAR.

**Table 1.** Simulation parameters.

Name	Geo-SAR	Leo-SAR
Orbital altitude (km)	35792	514
Inclination ( $^{\circ}$ )	60	97.42
Elevation angle ( $^{\circ}$ )	4.8	37.5
Wavelength (m)	0.24	0.03
Eccentricity	0.0011	
Longitude ascending node ( $^{\circ}$ )	0	
perigee argument ( $^{\circ}$ )	90	
$J_2$	$1082.63 \times 10^{-6}$	
$J_3$	$-2.5356 \times 10^{-6}$	
$J_4$	$-1.62336 \times 10^{-6}$	

### 5.1. Calculation Errors Caused by Orbital Perturbation

The error analysis shows that the satellite elliptical trajectory in two-body condition could be mainly influenced by  $J_2$ ,  $J_3$  and  $J_4$  perturbation terms. If the perturbation force is ignored, the 1–4th order Doppler parameters may have errors. Fig. 4 simulates the error in percentage of the four Doppler parameters of Geo-SAR and Leo-SAR.



**Figure 4.** Doppler parameter calculation error in percentage (Caused by ignoring the orbital perturbation).

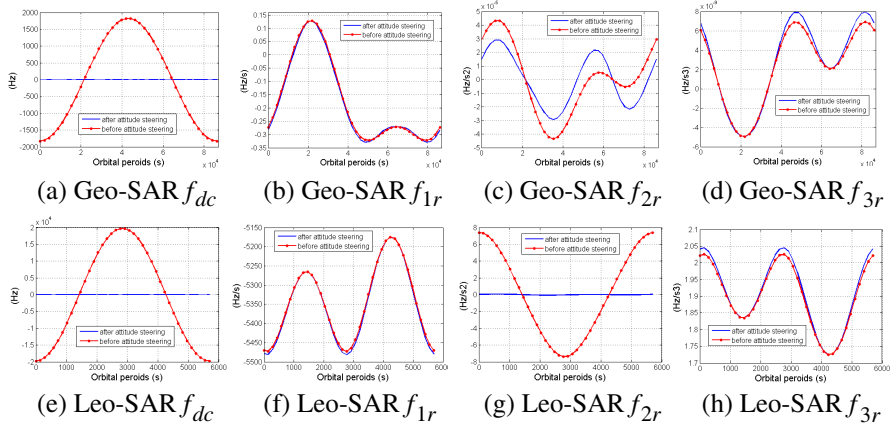
As shown in Fig. 4(a), the perturbation force mainly has an effect on Doppler FM rate ( $f_{1r}$ ) for Geo-SAR, and the largest error in percentage of  $f_{1r}$  is about 2.7%. Its influence on the other three Doppler parameters is not significant, however, small error in percentage below 0.7% can be observed. For Leo-SAR, the largest error in percentage of  $f_{3r}$  is about 9.8% and the errors of the other three Doppler parameters are all below 4.8%. Since the semi-major axis of Leo-SAR is around 1/6 of that of Geo-SAR, the perturbation force has effect on the six orbital elements tremendously for the former. As a result, it is reasonable that the calculation errors in percentage of Leo-SAR’s Doppler parameters are larger than those of Geo-SAR’s.

### 5.2. The Effect of Attitude Steering on Doppler Parameters

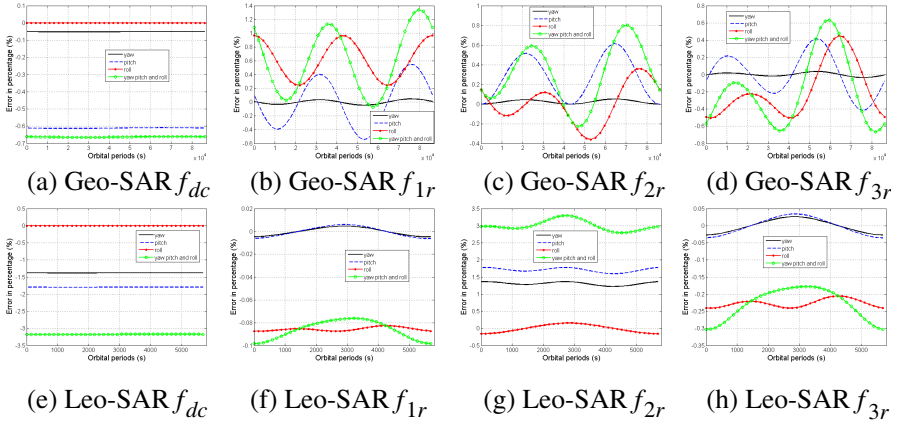
Since the beam pointing direction changes, the Doppler parameters may change significantly after attitude steering. The simulation results are depicted in Fig. 5.

In Fig. 5, it is obvious that the absolute values of Leo-SAR’s Doppler parameters are larger than those of Geo-SAR’s. In addition, the absolute values of Geo-SAR’s change more sharply with respect to time. Besides the Doppler parameter properties of Leo-SAR and Geo-SAR are significantly different at the following two aspects:

- The Doppler centroid properties of both systems are similar, but the characteristics of  $f_{1r}$ ,  $f_{2r}$  and  $f_{3r}$  are quite different.
- The attitude steering has an obvious effect on the values of  $f_{dc}$  and  $f_{2r}$  for both Leo-SAR and Geo-SAR. However,  $f_{1r}$  and  $f_{3r}$  are approximately independent of the beam pointing directions.



**Figure 5.** Doppler parameter properties comparison between Geo-SAR and LeoSAR.



**Figure 6.** Doppler parameter error in percentage (Caused by attitude control error).

### 5.3. Attitude Control Accuracy

Any mispointing of radar beam, caused by the attitude control error, may lead to a shift of the Doppler parameters. The beam pointing errors of  $0.05^\circ$  for yaw, pitch and roll angles are assumed in this paper. The effects of pointing errors on the 1–4th order Doppler parameters for Geo-SAR and LeoSAR are analyzed and simulated in Fig. 6.

According to the simulation results in Fig. 6, we can get the conclusions as follows:

- For a constant attitude control error, the error in percentage of Doppler centroid for Both Geo-SAR and Leo-SAR remains almost constant during the whole orbital periods. However, the errors of the other three Doppler parameters always change with time.
- All the errors of the three attitude angles could affect the Doppler parameters of the two spaceborne systems.
- The error of percentage of  $f_{dc}$  and  $f_{2r}$  for Leo-SAR are larger than those of Geo-SAR's, however, the pointing accuracy could have an obvious effect on the values of  $f_{1r}$  and  $f_{3r}$  for Geo-SAR.

## 6. CONCLUSION

It is necessary to obtain the accurate 1–4th order Doppler parameters for Geo-SAR. The calculation expressions of the Doppler parameters are derived based on an accurate geometry model. The  $J_2$ ,  $J_3$  and  $J_4$  orbital perturbation terms could have an obvious influence on the trajectory of satellite. If the orbital perturbation force is ignored, the four Doppler parameters could have errors for both Geo-SAR and Leo-SAR. The largest error in percentage of Geo-SAR's Doppler parameters is about 2.7%. However, this value of Leo-SAR is around 9.8% since its orbital height is low and the effect of perturbation force is obvious. The absolute values of Doppler parameters of Geo-SAR are quite small and change evidently with respect to time compared with Leo-SAR. Besides the properties of them except the Doppler centroid frequency are quite different between Geo-SAR and Leo-SAR. For both Geo-SAR and Leo-SAR,  $f_{1r}$  and  $f_{3r}$  are approximately not affected by attitude steering, but the values of  $f_{dc}$  and  $f_{2r}$ , on the other hand, are sensitive to the satellite attitude. Finally, the attitude control error could also have an effect on the Doppler parameters for the two spaceborne SAR systems.

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