The Axial Ratio of Planar Arrays with Random Element Errors

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Abstract—Characterizing the random errors at the elements of a phased array antenna leads to equations that estimate the associated performance degradation. The increase in sidelobe level and decrease in gain due to random errors is well established. This paper derives an expression that predicts the axial ratio degradation due to random errors in the circularly polarized elements of an array. In the case of small errors in an array of crossed dipoles, we found a simple expression for the axial ratio of the array under random errors at broadside.

1. INTRODUCTION

Circularly polarized antenna arrays have many practical applications in communications and radar systems. Satellite communication arrays rely on circular polarization due to Faraday rotation of the signal as it passes through the ionosphere. Meteorological radar arrays must have very high isolation between vertical and horizontal polarization in order to distinguish between different types of precipitation. Calibrating errors enhances the polarization isolation [1, 2].

Random array errors due to manufacturing tolerances, equipment aging, and temperature raise the sidelobe level and lower the gain of phased arrays. Traditionally, random errors in arrays have been analyzed to determine the impact on the sidelobe level. This paper looks at how random amplitude and phase errors change the axial ratio of a dual or circularly polarized planar array.

Ruze first analyzed the impact of aperture errors on antenna patterns [3]. Random errors in an array are considered statistically independent from element to element. In this paper, position errors and random element failures are ignored, but amplitude and phase errors are not. These errors produce very broad array patterns that are superimposed on the no-error patterns resulting in "filled-in" nulls and the increased gain of low sidelobes [4, 5].

To the knowledge of the authors the analysis of the impact of array errors on the polarization purity of an array has not been done before (besides preliminary results in [6]). The goal of this paper is to model a circularly polarized planar array by an array of crossed dipoles with errors on the feeding currents in order to quantify the impact of array errors on the polarization.

In Section 2 we present the crossed dipoles array model for the planar array, and we define the errors in the feeding currents. In Section 3 we derive the approximate expression of the Axial Ratio (AR) as a result of the errors. In Section 4 we validate the resulting equation against the original model and observe the effect of the errors with random trial simulations.

2. CROSSED DIPOLE ARRAY MODEL

In this section we present a model of a circularly polarized planar array of crossed dipoles with random amplitude and phase errors. In order to find the AR of the crossed dipoles array, we calculate the

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orthogonal linear electric fields of the array. We restrict the analysis to right handed polarized crossed dipoles in this paper, but the results apply to left handed polarized arrays as well because of the symmetry of the problem. If the two crossed dipoles have the same length, with one along the x-axis and the other along the y-axis [7]:

$$E_{\theta} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \left[I_h \cos(\theta) \cos(\phi) + I_v e^{j\frac{\pi}{2}} \cos(\theta) \sin(\phi) \right]$$
(1)

$$E_{\phi} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \left[-I_h \sin(\phi) + I_v e^{j\frac{\pi}{2}} \cos(\phi) \right]$$
(2)

where ω is the angular frequency; μ is the space magnetic permeability; k is the wave number; r is the distance from the antenna to the observation point; I_h is the current on the horizontally oriented dipole; I_v is the current on the vertically oriented dipole; L is the length of the dipoles; θ is measured from the z-axis; and ϕ is measured from the x-axis.

If we neglect the mutual coupling between the different elements but feed each of the crossed dipoles with a different current, then the electric field of an array of N crossed dipoles is given by

$$E_{\theta} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \sum_{i=1}^{N} \left\{ e^{-j\delta_{pos_i}} \left[I_{hi} \cos(\theta) \cos(\phi) + I_{vi} e^{j\frac{\pi}{2}} \cos(\theta) \sin(\phi) \right] \right\}$$
(3)

$$E_{\phi} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \sum_{i=1}^{N} \left\{ e^{-j\delta_{pos_i}} \left[-I_{hi} \sin(\phi) + I_{vi} e^{j\frac{\pi}{2}} \cos(\phi) \right] \right\}$$
(4)

$$\delta_{pos_i} = \frac{2\pi}{\lambda} \Big\{ \big[\sin(\theta) \cos(\phi) - \sin(\theta_s) \cos(\phi_s) \big] x_i + \big[\sin(\theta) \sin(\phi) - \sin(\theta_s) \sin(\phi_s) \big] y_i \Big\}$$
(5)

where $[x_i, y_i]$ is the position of the *i*th crossed dipoles element. Errors are assumed to be amplitude and phase deviations from the desired current.

$$I_{hi} = I \left(1 + \delta I_{hi} \right) e^{-j\phi_{hi}} \tag{6}$$

$$I_{vi} = I \left(1 + \delta I_{vi} \right) e^{-j\phi_{vi}} \tag{7}$$

where δI_{hi} , δI_{vi} , $\delta \phi_{hi}$, and $\delta \phi_{vi}$ are, respectively: the horizontal dipole normalized amplitude error, the vertical dipole normalized amplitude error, the horizontal dipole phase error, and the vertical dipole phase error for the *i*-th crossed dipoles element.

3. AXIAL RATIO FORMULA FOR AN ARRAY WITH RANDOM ELEMENT ERRORS

This section presents the derivation of the AR expression for the crossed dipoles array with errors. By means of a simple normalization, observing the broadside radiation without steering $([\theta, \phi] = [\theta_s, \phi_s] = [0, 0])$ and using Equations (6) and (7), (3) and (4) reduces to the following equations:

$$E_{\theta} = \sum_{i=1}^{N} \left(1 + \delta I_{hi} \right) e^{-j\delta\phi_{hi}} \tag{8}$$

$$E_{\phi} = e^{j\frac{\pi}{2}} \sum_{i=1}^{N} (1 + \delta I_{vi}) e^{-j\delta\phi_{vi}}$$
(9)

At this point we consider the errors to be small and make the Taylor expansion $e^x = 1 + x + O(x^2)$, as well as neglect the terms $\delta I_{hi} \delta \phi_{hi}$ and $\delta I_{vi} \delta \phi_{vi}$. Next, decomposing the fields into right and left handed polarization components using (A1) and (A2) (see Appendix), results in:

$$E_{rhp} = \frac{1}{\sqrt{2}} \sum_{i=1}^{N} \left(2 + \delta I_{hi} + \delta I_{vi} - j\delta\phi_{hi} - j\delta\phi_{vi} \right)$$
(10)

$$E_{lhp} = \frac{1}{\sqrt{2}} \sum_{i=1}^{N} \left(\delta I_{hi} - \delta I_{vi} - j \delta \phi_{hi} + j \delta \phi_{vi} \right) \tag{11}$$

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These equations simplify to:

$$E_{rhp} = \frac{1}{\sqrt{2}} \left(2N + \delta I_+ - j\delta\phi_+ \right) \tag{12}$$

$$E_{lhp} = \frac{1}{\sqrt{2}} \left(\delta I_{-} - j \delta \phi_{-} \right) \tag{13}$$

where:

$$\delta I_{-} = \sum_{i=1}^{N} \left(\delta I_{hi} - \delta I_{vi} \right) \tag{14}$$

$$\delta\phi_{-} = \sum_{i=1}^{N} \left(\delta\phi_{hi} - \delta\phi_{vi}\right) \tag{15}$$

$$\delta I_{+} = \sum_{i=1}^{N} \left(\delta I_{hi} + \delta I_{vi} \right) \tag{16}$$

$$\delta\phi_{+} = \sum_{i=1}^{N} \left(\delta\phi_{hi} + \delta\phi_{vi}\right) \tag{17}$$

Finally, by using Equations (A5) (see Appendix A), (12) and (13) we get:

$$AR = \frac{20}{N\ln(10)} \sqrt{\frac{\delta I_{-}^2 + \delta \phi_{-}^2}{\left[\left(1 + \frac{\delta I_{+}}{2N}\right)^2 + \left(\frac{\delta \phi_{+}}{2N}\right)^2\right]}}$$
(18)

Assuming that the terms $\frac{\delta I_+}{2N}$ and $\frac{\delta \phi_+}{2N}$ are small enough to neglect, then we get:

$$AR = \frac{20}{N\ln(10)}\sqrt{\delta I_{-}^{2} + \delta \phi_{-}^{2}}$$
(19)

The AR of the array depends on the difference between the amplitudes and phases of the horizontal and vertical dipole currents. Random errors at the elements average, so Equation (19) is proportional to $\frac{1}{N}$.

4. RESULTS

In this section, results at boresight and scan angle are presented.

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4.1. Axial Ratio at Boresight

In order to validate Equation (19), we calculated the AR of one crossed dipole element for 10^6 random amplitude and phase errors. The errors are zero-mean, normally distributed with standard deviations of $\sigma_{\phi_h} = \sigma_{\phi_v} = 0.2$ (11.46°) and $\sigma_{I_h} = \sigma_{I_v} = 0.1$ (10%). We calculated the AR using Equation (19), then created the histograms in Figure 1 and calculated the figures of merit shown in Equations (20), (21), (22), and (23). We used the formula $(2n)^{1/3}$, with *n* the sample size (10⁶), to calculate the number of bins in the histograms [8]. To calculate the bin width for the histogram we considered the range of values to be contained between 0 dB and 10 dB, which gives a bin width of 0.08 dB.

$$|AR - AR_{approx}|_{99\%} = 0.68 \,\mathrm{dB} \tag{20}$$

$$|AR - AR_{approx}|_{95\%} = 0.32 \,\mathrm{dB}$$
 (21)

$$\frac{1}{N} \sum_{i=1}^{N} |AR_i - AR_{approx}i| = 0.088 \,\mathrm{dB}$$
(22)



Figure 1. Histogram of the random trial simulations of the AR calculations for a single crossed dipole element at broadside. (a) With approximations in (19), (b) without approximations. In dashed red we have the AR without errors (0 dB). The bins in the histogram have a 0.08 dB width.

$$\frac{1}{N}\sum_{i=1}^{N} \left| \frac{AR_i - AR_{approx}i}{AR_i} \right| = 3\%$$
(23)

This example demonstrates that the approximate equations derived in the last section accurately predict the change in AR due to small random errors at the elements.

In the case where all the errors are independent identically distributed Gaussian with zero mean, Equation (19) follows a Rayleigh distribution. By looking at the histograms we can see that even when the errors have different standard deviations they present a Rayleigh-like behaviour.

In order to observe the effect of having more than one element, we made the same test for a small uniform 5×5 planar array with a spacing of $\lambda/2$ (Figure 2). We can observe that, as it can be predicted from Equation (19), the more elements we add to the array, the smaller the variance on the results, which get closer to the ideal case without errors.



Figure 2. Histogram of the random trial simulations of the AR calculations without approximations for 25 crossed dipoles elements looking at broadside. In dashed red we have the AR without errors (0 dB). The bins in the histogram have a 0.08 dB width.

4.2. Axial Ratio at Scan Angle

Scanning the main beam away from broadside makes the calculations for the AR much more complicated. However, it is simple to make a random trial simulation of the AR at the steering direction. Using Equations (3), (4), and (5) and making $[\theta, \phi] = [\theta_s, \phi_s]$ we get:

$$E_{\theta} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \sum_{i=1}^{N} \left[I_{hi} \cos(\theta_s) \cos(\phi_s) + I_{vi} e^{-j\frac{\pi}{2}} \cos(\theta_s) \sin(\phi_s) \right]$$
(24)

$$E_{\phi} = -j \frac{\omega \mu L e^{-jkr}}{4\pi r} \sum_{i=1}^{N} \left[-I_{hi} \sin(\phi_s) + I_{vi} e^{-j\frac{\pi}{2}} \cos(\phi_s) \right]$$
(25)

Note that in the steering direction, the radiated fields are no longer a function of the element positions, since the electronic phase applied to each element compensates the phase due to the position of the element in the steering direction.

For this simulation, we will consider first a single element looking at the steering direction (Figure 3), then a small uniform 5×5 planar array with a spacing of $\lambda/2$ (Figure 4). The errors are independent Gaussian distributed, with the amplitude errors having a standard deviation $\sigma_{I_h} = \sigma_{I_v} = 0.1$ (10%) and the phase errors $\sigma_{\phi_h} = \sigma_{\phi_v} = 0.2$ (11.46°). For the simulation we will consider a steering of $[\theta_s, \phi_s] = [30^\circ, 0^\circ]$, which give an AR of 1.25 dB for an array without errors.



Figure 3. Histogram of the random trial simulations of the AR calculations without approximations for a single crossed dipoles element at 30° . In dashed red we have the AR without errors (1.25 dB). The bins in the histogram have a 0.08 dB width.



Figure 4. Histogram of the random trial simulations of the AR calculations without approximations for 25 crossed dipoles elements scanned to 30° . In dashed red we have the AR without errors (1.25 dB). The bins in the histogram have a 0.08 dB width

Figure 3 shows that the AR of the isolated element at the steering direction degrades when errors are considered. Comparing the broadside case in (Figure 1) with the scanned case in Figure 3, we observe that the AR degrades off-broadside.

In Figure 4, we observe that the AR from the random trial simulations is nearly the same as the estimate calculated using errorless feeding currents. The random trial simulations get closer to the errorless case as the number of elements in the array increases.

5. CONCLUSION

By analyzing the equations for a radiating element composed of two crossed dipoles with errors in amplitude and phase, we were able to find a simple formulation for the resulting AR. This estimate depends on the difference between the amplitudes and phases of the horizontal and vertical dipole currents. Random trial simulations validate the approximate formulas for broadside arrays and are used to explore scanned arrays.

We observe that considering the feeding errors to be independent identically distributed Gaussian variables, the expression in Equation (19) leads to a Rayleigh probability density function. Equation (19) also predicts that, as the number of elements (N) in the array gets larger, random errors have less impact on the AR of the array due to averaging, which is verified in Figure 2 and Figure 4.

APPENDIX A. AXIAL RATIO EXPRESSION

Having the linear polarized fields (E_{θ} and E_{ϕ}), we can calculate the right handed and left hand polarized electric fields $(E_{rhp} \text{ and } E_{lhp})$ using the following equations:

$$E_{rhp} = \frac{1}{\sqrt{2}} \left(E_{\theta} - j E_{\phi} \right) \tag{A1}$$

$$E_{lhp} = \frac{1}{\sqrt{2}} \left(E_{\theta} + j E_{\phi} \right) \tag{A2}$$

We can then calculate the cross polarization ratio and find the AR with the following equations:

$$xpolr^2 = \frac{|E_{rhp}|^2}{|E_{lhp}|^2}$$
 (A3)

$$AR = 20\log_{10}\left(\frac{xpolr+1}{xpolr-1}\right), \quad xpolr > 1$$
(A4)

which means that $|E_{rhp}| > |E_{lhp}|$ (a reasonable assumption for a right handed polarized antenna). By taking the Laurent series of $\ln(\frac{x+1}{x-1})$, as $x \to \infty$, we get $\frac{2}{x} + O(\frac{1}{x^3})$, which gives an approximation for the AR expression, that becomes the following by combining Equations (A3) and (A4):

$$AR = \frac{40}{\ln(10)} \sqrt{\frac{|E_{lhp}|^2}{|E_{rhp}|^2}}$$
(A5)

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